

Lecture 5: Classical Mechanics

Particle motion in classical mechanics: displacement of particle $\underline{x}_p(t)$ evolves according to Newton's 2nd Law:

Trajectory Eqn: $m \ddot{\underline{x}}_p(t) = \underline{F}(t, \underline{x}_p(t))$

Initial conditions: $\underline{x}_p(0) = \underline{x}_0$, $\dot{\underline{x}}_p(0) = \underline{v}_0$

Laplace's Demon: given ICs, can predict $\underline{x}_p(t)$ for all time.

- this perspective upended by both the probabilism of QM, and the limits on predictability imposed by chaos.

Chaotic Dynamics: for sufficiently complex systems, uncertainty in ICs ($\underline{x}_p(0) = \underline{x}_0 + \underline{\epsilon}_p$, $\dot{\underline{x}}_p(0) = \underline{v}_0 + \underline{\epsilon}_v$) are amplified exponentially with time
 \Rightarrow loss of predictability

Ex. 1 the Butterfly effect

Ex. 2 in orbital pilot-wave dynamics, chaos characterized by intermittent switching between unstable periodic states

Hereditary Systems: classical systems whose evolution depends explicitly on their past

\Rightarrow a.k.a. time-delay, non-Markovian, or temporally nonlocal systems

- to predict the future of a hereditary system, one needs to know its history
- originally considered by Brillouin, Boltzmann, Volterra (~1910)

Eg. 1 For elastic solids, the state of stress depends on the system's history of strain

Eg. 2 For an electric charge influenced by its own field, the force on the charge may depend on its history

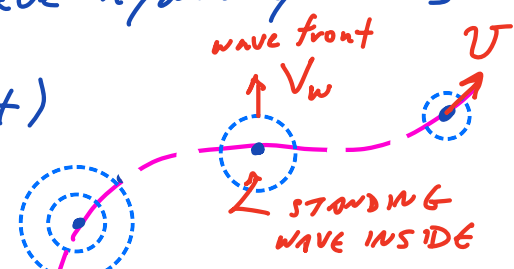
Eg. 3 For particle motion in a fluid, the stress field will generally depend on the particle history
 e.g. a fish swimming in a circle [SLIDE]

Trajectory equation : $m \ddot{\underline{x}}_p(\tau) = \underline{F}(\tau, \underline{x}_p(t < \tau))$

• described by delay differential equations: the particle acceleration depends on particle history

Eg. the Stroboscopic Model of pilot-wave hydrodynamics

$$m \ddot{\underline{x}}_p = - \underbrace{D \dot{\underline{x}}_p}_{\text{DRAG}} - mg \underbrace{\nabla h(\underline{x}_p, t)}_{\text{WAVE FORCE}}$$



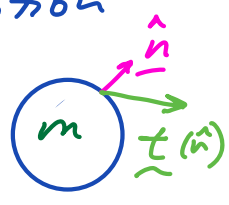
where $\nabla h(\underline{x}_p, t) \sim \int_{-\infty}^t \underline{K}(\underline{x}_p(t) - \underline{x}_p(s)) ds$
 NB: $V_w \gg U$

Note: the system appears to be hereditary only when the trajectory equation is considered in isolation

• a complete formulation requires description of both the particle motion and the associated fluid motion

FLUID MECHANICS

Trajectory equation for a mass moving in a fluid : $m \ddot{\underline{x}}_p = \int_S \underbrace{\hat{n} \cdot \underline{T}}_{\text{stress vector}} ds$



where $\underline{T}(p, \underline{u})$ is the stress tensor, and the fluid pressure p and velocity \underline{u} evolve according to Navier-Stokes equations.

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Throughout the course, we restrict our attention to incompressible, Newtonian fluids with constant density ρ and dynamic viscosity $\mu = \rho \nu$.

Note: the kinematic viscosity $\nu = \frac{\mu}{\rho}$ has units $\left[\frac{L^2}{T}\right]$ so prescribes the rate of diffusion of momentum and vorticity within the flow

For such fluids, the Navier-Stokes equations are a set of 4 equations in 4 unknown field variables, specifically the pressure $p(\underline{x}, t)$ and the 3 components of the velocity $\underline{u}(\underline{x}, t)$. They follow from the application of the classical concepts of conservation of mass, momentum and angular momentum to an infinitesimal fluid element.

ASSUMPTION
Continuum Hypothesis: the fluid is still smooth at this scale $\delta \gtrsim 10$ fluid molecules, so the graininess associated with molecular structure need not be considered.

Conservation of Mass: $\underline{\nabla} \cdot \underline{u} = 0$

Conservation of Momentum:

$$\rho \left(\underbrace{\frac{\partial \underline{u}}{\partial t}}_{\text{INERTIA}} + \underbrace{\underline{u} \cdot \underline{\nabla} \underline{u}}_{\text{APPLIED FORCE}} \right) = \underbrace{\underline{f}}_{\text{APPLIED FORCE}} + \underbrace{\underline{\nabla} \cdot \underline{T}}_{\text{FORCE ASSOCIATED WITH STATE OF STRESS}}$$

is Cauchy's Momentum Eqn

Conservation of Angular Momentum

Stress tensor symmetric: $\underline{\underline{T}} = \underline{\underline{T}}^T$

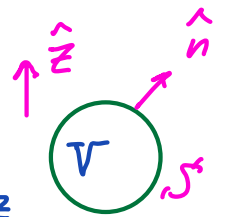
Express stress tensor: $\underline{\underline{T}} = -p \underline{\underline{I}} + \underline{\underline{\tau}}(\underline{u})$

DEVIATORIC STRESS TENSOR
STATIC DYNAMIC

Static component consistent with our notion of pressure.

Force on a submerged body in a static fluid

$$\underline{F}_h = \int_S \hat{n} \cdot \underline{\underline{T}} dS = - \int_S p \hat{n} dS$$



eg. with hydrostatic pressure $p(z) = p_0 - \rho g z$

$$\underline{F}_h = - \int_S (p_0 - \rho g z) \hat{n} dS = \rho g \int_V \nabla z dV = \rho g V \hat{z}$$

GEN DIV THM

BUOYANCY FORCE = weight of displaced fluid

⇒ Archimedes Principle

We now need only deduce the form of the deviatoric stress tensor $\underline{\underline{\tau}}(\underline{u})$.

Since $\underline{\underline{T}} = -p \underline{\underline{I}} + \underline{\underline{\tau}}$ is symmetric, so too must be $\underline{\underline{\tau}}$.

Moreover, $\underline{\underline{\tau}}$ cannot depend on translational or rotational velocity, or it would be frame dependent. We thus expect

$\underline{\underline{\tau}}$ to depend on the symmetric rate of strain tensor

$$\underline{\underline{E}} = \frac{1}{2} \left[\underline{\underline{\nabla u}} + (\underline{\underline{\nabla u}})^T \right]$$

Newtonian Fluids

(1) Relation between $\underline{\underline{\tau}}$ and $\underline{\underline{E}}$ is local in space + time.

Counterexample: fluids with 'memory', for which $\underline{\underline{T}}$ depends on the history of strain e.g. elastic fluids, polymers

(2) Relation is linear. The most general such relationship may be expressed as $\underline{\underline{T}} = \underline{\underline{A}} : \underline{\underline{E}}$ where $\underline{\underline{A}}$ is a 4th rank constant tensor.

(3) Fluid is isotropic: there is no preferred orientation in space $\Rightarrow \underline{\underline{A}}$ must be an isotropic 4th rank tensor

For the incompressible Newtonian fluids of interest, this relation assumes the simple form

$$\underline{\underline{T}} = 2\mu \underline{\underline{E}} \quad \text{where } \mu \text{ is the fluid viscosity}$$

We may thus express the stress tensor:

$$\underline{\underline{T}} = -p \underline{\underline{I}} + 2\mu \underline{\underline{E}}$$

Substituting into Cauchy's Momentum Eqn and noting

$$\underline{\nabla} \cdot \underline{\underline{T}} = -\underline{\nabla} p + 2\mu \underline{\nabla} \cdot \underline{\underline{E}} = -\underline{\nabla} p + \mu \nabla^2 \underline{u}$$

then yields

Navier-Stokes Eqns for an incompressible Newtonian fluid with constant ρ, μ :

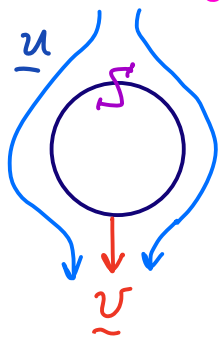
$$\rho \left(\frac{d\underline{u}}{dt} + \underline{u} \cdot \underline{\nabla} \underline{u} \right) = -\underline{\nabla} p + \mu \nabla^2 \underline{u} + \underline{f}$$

$$\underline{\nabla} \cdot \underline{u} = 0$$

4 eqns in 4 unknowns (\underline{u}, p) must be solved subject to appropriate initial and boundary conditions.

Fluid - Solid BCs : no-slip

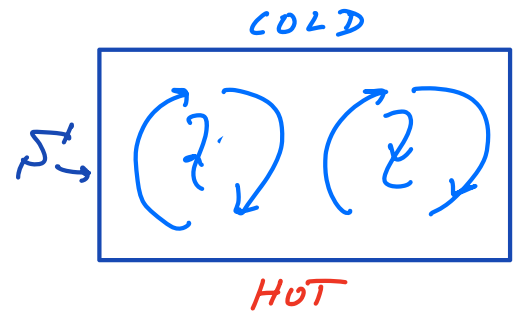
$$\Rightarrow \underline{u} = \underline{v}_{\text{SOLID}}$$



Eg. 1 Falling sphere : $\underline{u} = \underline{v}$ on S

Eg. 2 Convection in a box :

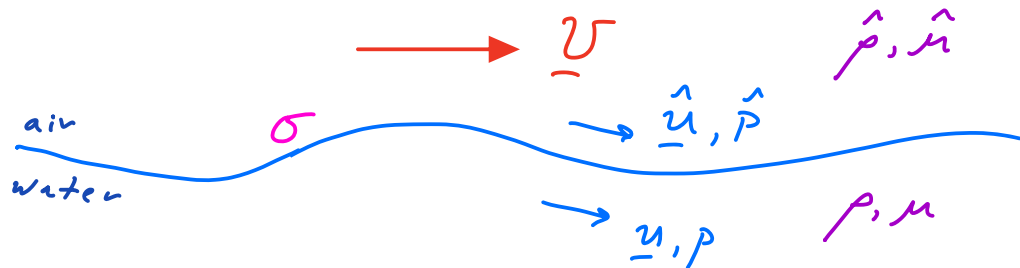
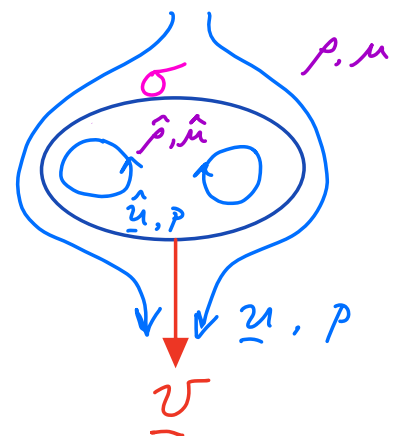
$$\underline{u} = 0 \text{ on } S'$$



In this course, we are concerned with flows dominated by interfacial effects

e.g. drop motion

e.g. water waves



Note : these interfaces are free to move ; thus, this of problems are known as **FREE BOUNDARY PROBLEMS**

Continuity of Velocity at an interface requires that $\underline{u} = \underline{\hat{u}}$.

We shall proceed by developing the interfacial stress boundary conditions.

But first, what is an interface, and what is surface tension?

⇒ SLIDES